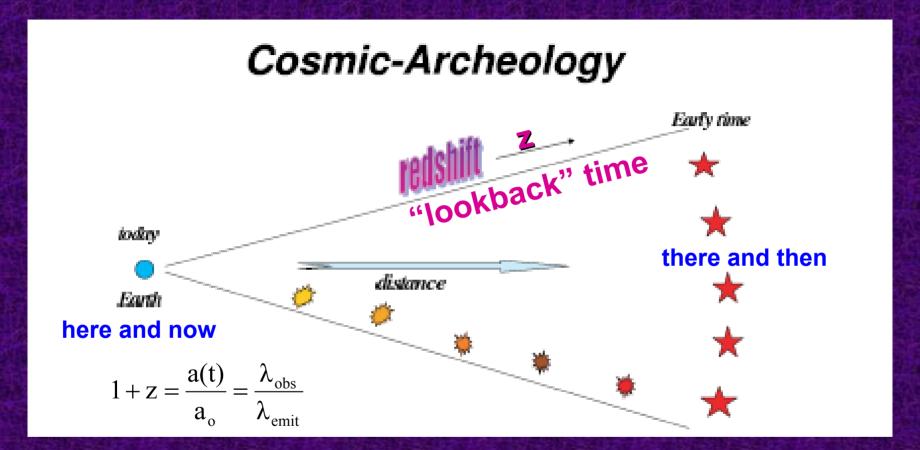
### Quasars and the Birth & Evolution of Galaxies

Kirk Korista

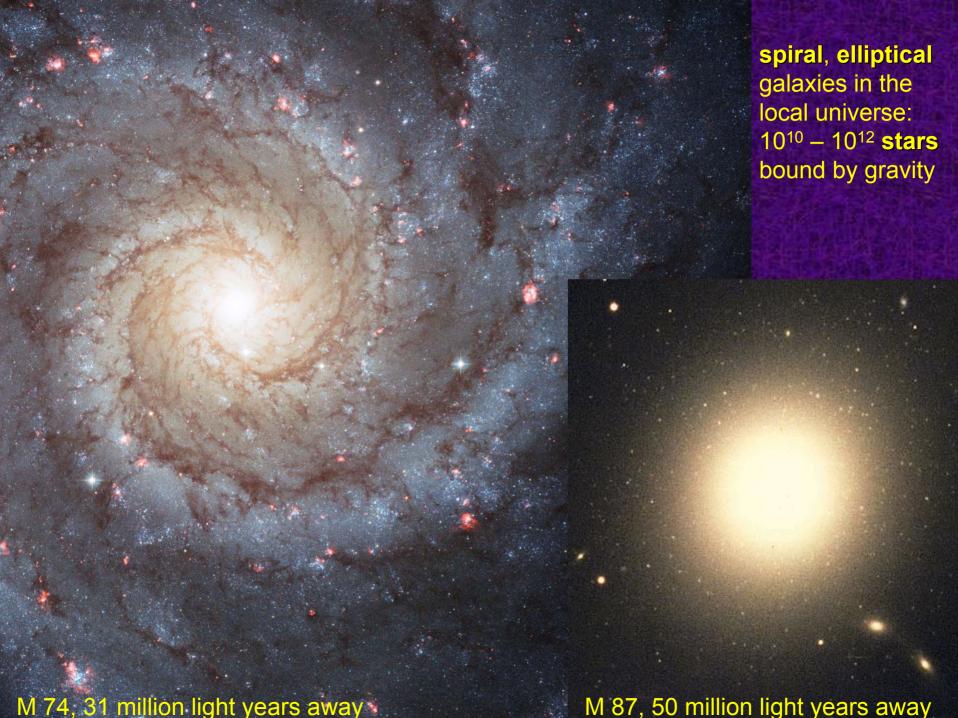
Associate Professor of Astronomy

Western Michigan University

#### Astronomers are "cosmic archeologists"



"Cosmology is like archeology. The deeper one looks the more ancient is the layer that is revealed, owing to the finite propagation speed of light." (Avi Loeb, 2006)



### **1963**-1983: Quasars were mysterious ultra-luminous points of light foreground normal galaxies Milky Way Galaxy star in the extreme foreground quasar

## Now: Quasars are the ultra-luminous "active" nuclei of massive galaxies...

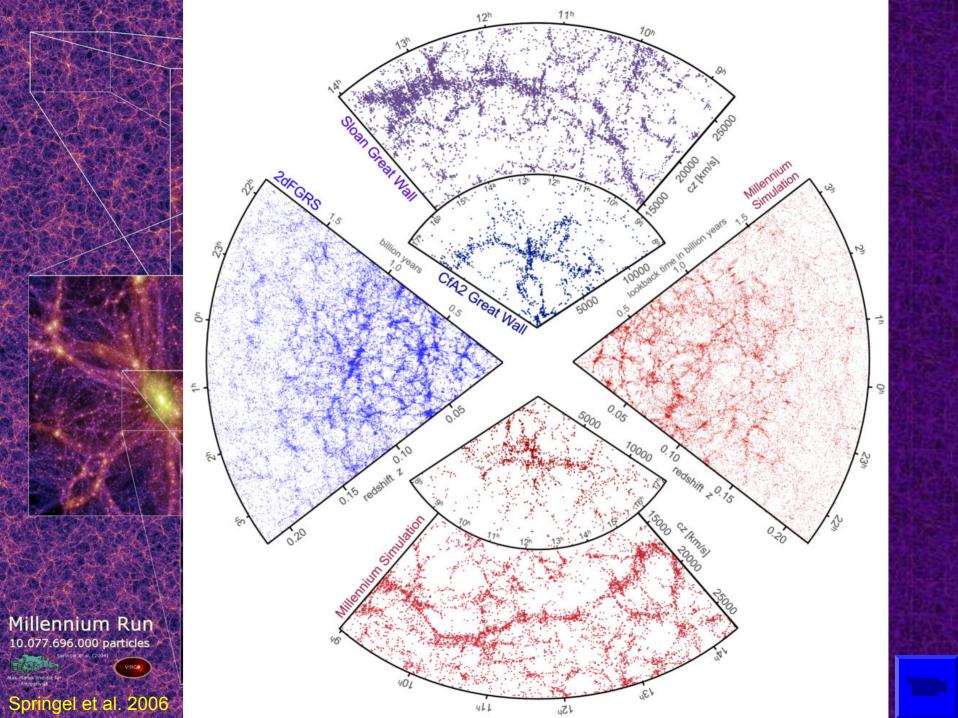
Quasar 3C 273, an "<u>AGN</u>"

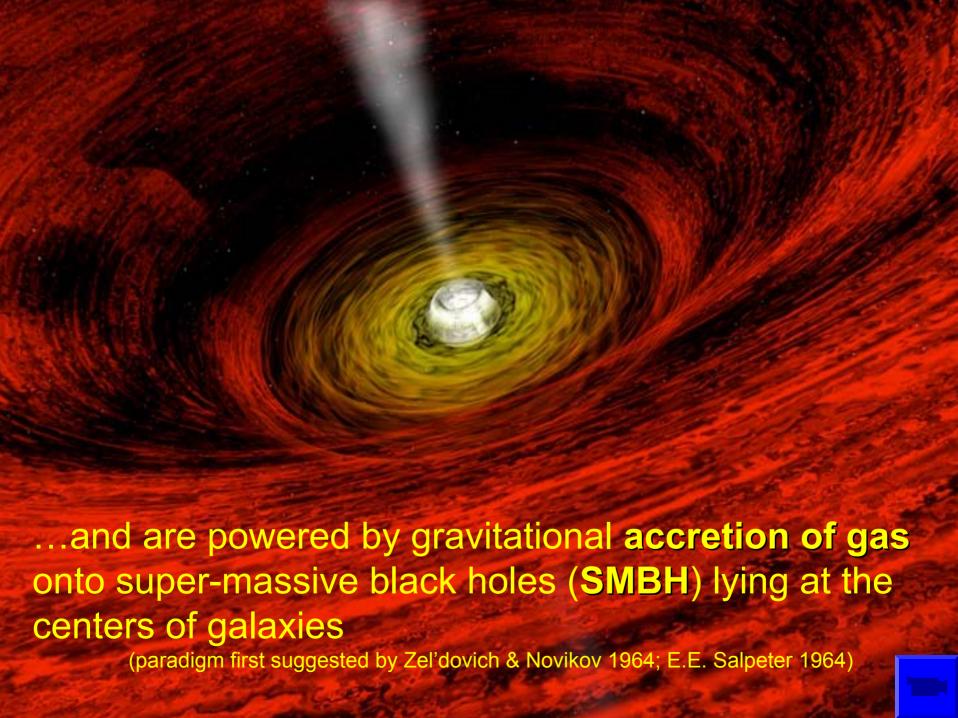
2 billion light years away, yet you can see this quasar with a backyard telescope! WFPC2 Host galaxy of 3C 273

ACS/HRC

NASA, A. Martel (JHU), the ACS Science Team, J. Bahcall (IAS) and ESA

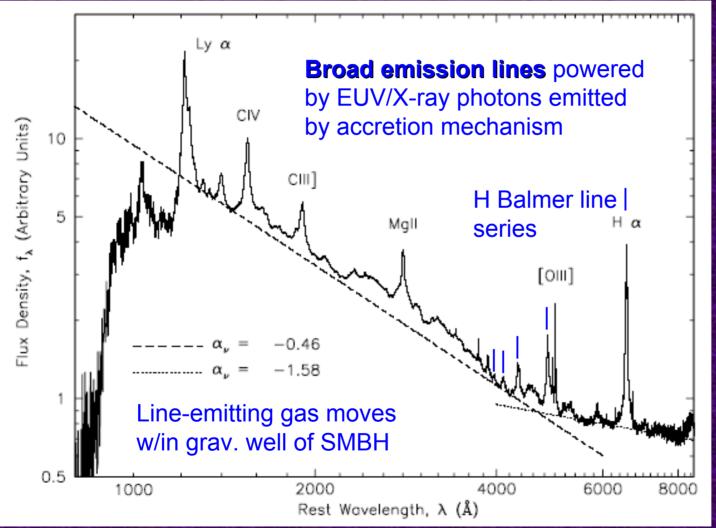
STScI-PRC03-03





#### A "generic" UV-optical spectrum of a Quasar

(a composite spectrum of ~2000 individual quasar spectra)



Heavy element emission lines observed to highest redshift quasars with  $z \sim 6.5$ :

universe just ~ 900 Myr old, so significant star formation episode must have occurred before then!

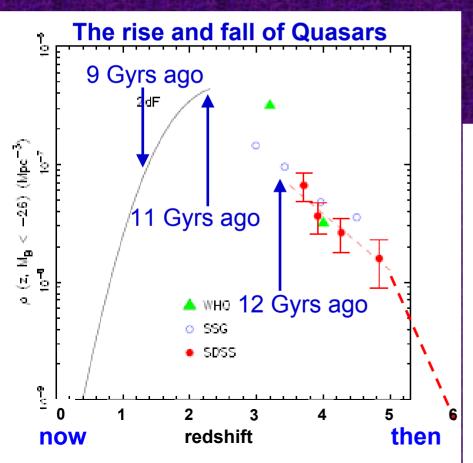
#### "Monster" Black Holes in Galactic Nuclei

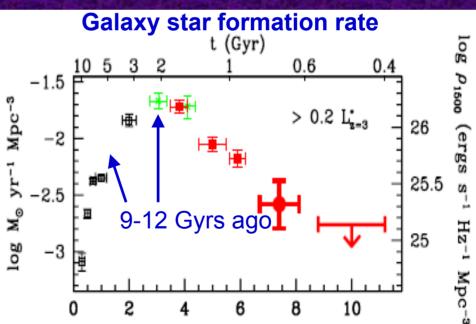
- With masses M = 10<sup>8</sup> several × 10<sup>9</sup> M<sub>sun</sub>, they're called super-massive black holes (aka "monsters")
- Accretion process releases ~10% of the rest mass of the accreted gas within region only a bit larger than solar system!
- $\rightarrow$   $L_{quasar} \approx 10^{12} 10^{15}$   $L_{sun},$  limited by max rate of accretion that scales with  $\mbox{\it M}$ 
  - (a lá Eddington's radiation pressure limit): L<sub>crit</sub> ~ L<sub>Edd</sub> ≈ 1.5 x 10<sup>39</sup> Watts · (M / 10<sup>8</sup> M<sub>sun</sub>)
- Luminous quasars are now extinct
  - nearest is 2 billion light years distant
  - low-power (10<sup>9</sup> 10<sup>12</sup> L<sub>sun</sub>) Active Galactic Nuclei ("AGN") sparsely populate *local universe*

# Quasars are interesting brutes of the cosmos, but so what??

# Rise and Fall of Quasar population mimics galaxy star formation rate through cosmic time

now

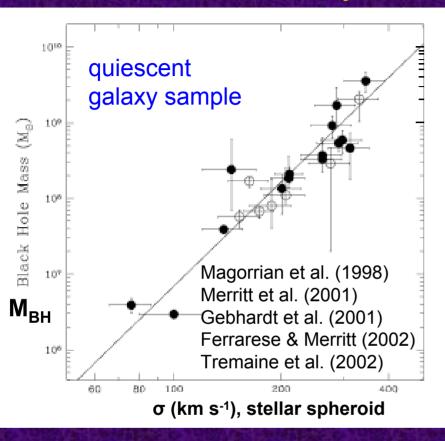


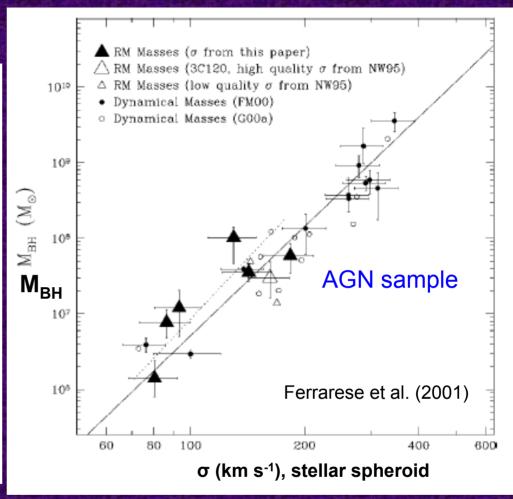


then

Bouwens et al. (2008)

#### And another curiosity...





All big galaxies (>  $10^{10}$  M<sub>sun</sub>) have supermassive black holes in their nuclei, AND < M<sub>BH</sub> / M<sub>stellar spheroid</sub>>  $\approx 0.002$  (!) (Ferrarese et al. 2006)

(→ physical size scale ratio is at least ~500, and so are dynamically disconnected)

#### And yet another...

...also scale with M<sub>BH</sub> →

Broad emission line intensity ratios that scale with gas **metallicity**Z ≡ Ab(heavy elements) / Ab(H)...

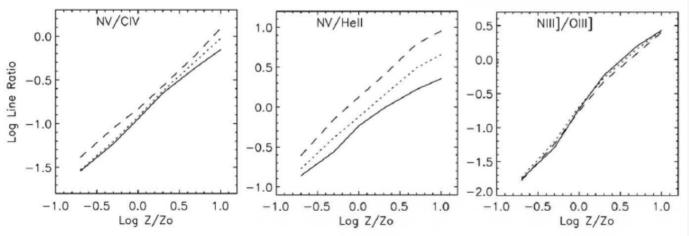
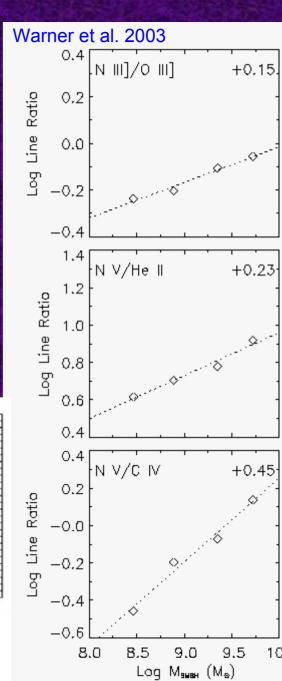
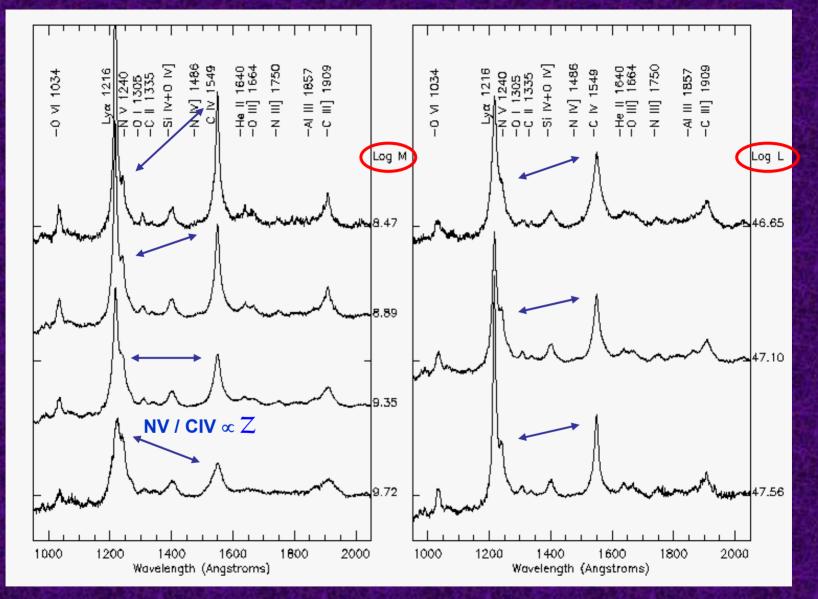


Figure 1. Theoretical BEL ratios vs. metallicity for three different ionizing spectral shapes (solid, dashed, dotted lines) in the LOC model (Hamann et al. 2002).





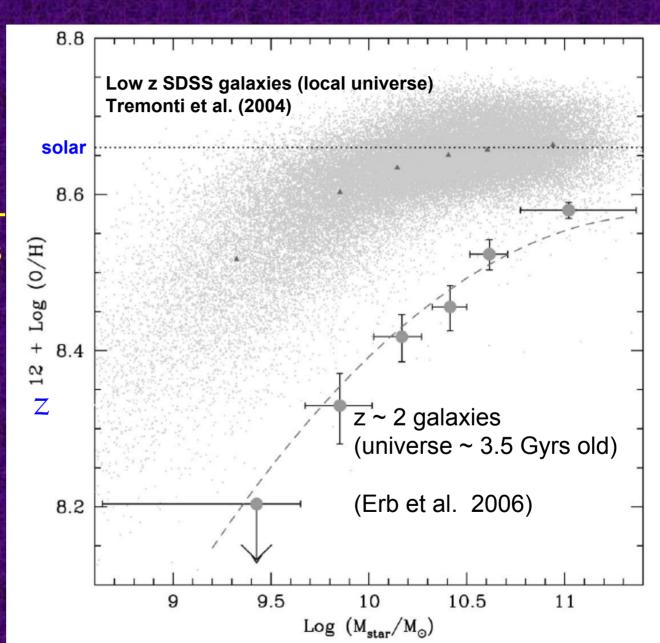
#### $M_{\rm BH} \approx 10^9 \, \rm M_{\rm sun}$



Composite quasar spectra for fixed L and  $M_{BH} \rightarrow$  an underlying trend is found between BH mass and Z (Warner et al. 2006)

that mirrors the well-known **galaxy** mass – metallicity relationship

(measured on galactic size scales)



#### Summary of findings so far...

- Quasars found in galactic nuclei, powered by accretion of matter onto super-massive black holes
- rise/fall of Quasar population mimics galaxy star formation rate through cosmic time
- <M<sub>BH</sub> / M<sub>stellar spheroid</sub> $> \approx 0.002$
- $M_{smbh}$ ,  $M_{stellar\ bulge}$  found independently to scale with gas metallicity Z on enormously different size scales

What's the connecting thread?

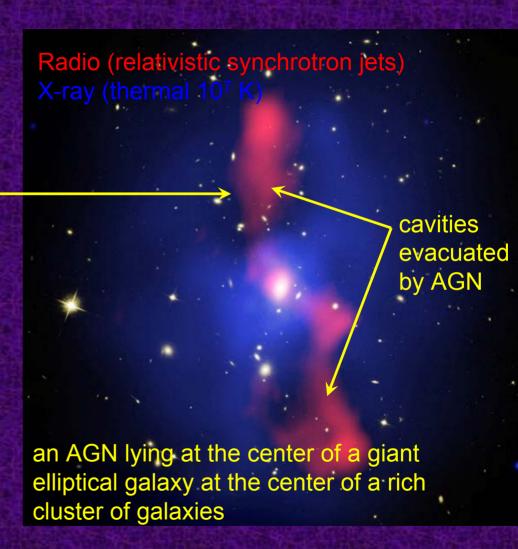
#### The quasar – galaxy formation connection

Silk & Rees 1998, Kauffman & Haehnelt 2000; Volonteri, Haardt, & Madau 2003; Granato et al. 2004

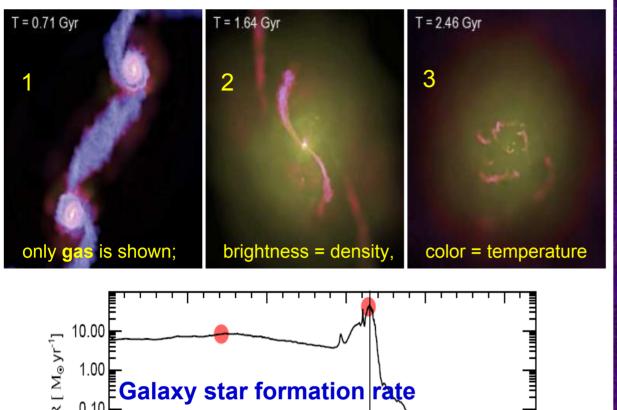
- Collisions between gas rich proto-galaxies drive gas toward and into the galactic nuclei, sparking prodigious star formation and feeding the monster(s)...
- ...until the monster becomes massive enough thus radiatively powerful enough to
  - heat and/or blow remaining gas out of the galactic well,
  - → throttling star formation and ultimately starving itself to death.
- Denser, more massive regions evolve more rapidly...
  - give birth to more stars (→ more massive galaxies)
  - are able to reach higher gas & star metallicities
  - have more gas available to feed SMBH
  - require that SMBH grows to greater mass before becoming powerful enough to blow out remaining gas from galaxy's deeper gravity well

#### Self-regulated growth via "AGN Feedback"

- radiatively powered winds
- photo-ionization heating of galactic environment → over-pressurization
- relativistic matter jets -
- various shock/pressure waves
- plus nuclear starbursts (supernovae, super winds)
  - but stars can't do it alone!
- The cosmic equivalent of a "2-liter Coke™ Chug"



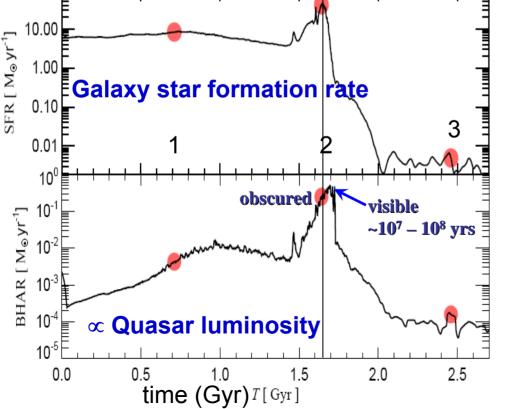
Galaxy cluster MS0735 (McNamara et al. 2006)

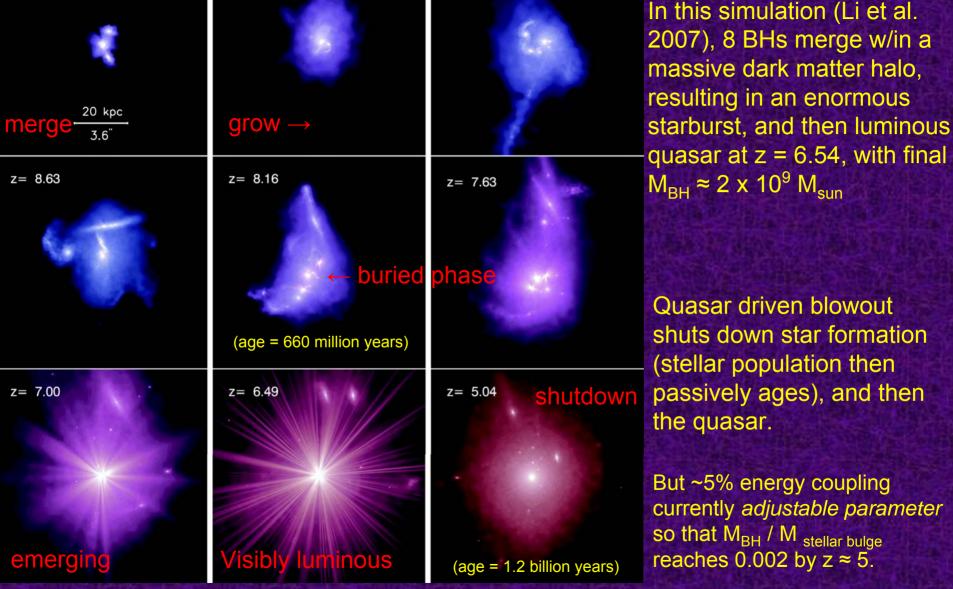


gas rich disk 'galaxies' merge in early universe...

- 1. After first pass: gas driven into nuclei, monsters grow; star formation rate high; gas metallicity grows...
- 2. Major merger completes with enormous amounts of gas driven into center, star form. rate zooms, luminous quasar turns on
- 3. Quasar "blowout" throttles star formation, *then* starves itself

Di Matteo et al. 2004; Hopkins et al. 2005; Springel et al. 2006





z = 9.17

intensity = stellar density; color = specific star formation rate; 'artificial' rays indicate quasar luminosity

z = 10.32

z = 12.75

(age = 360 million years)

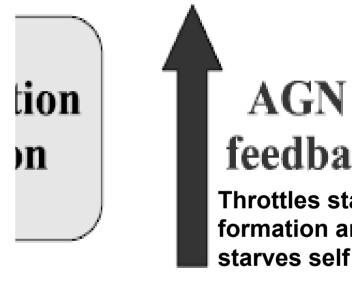
adapted from Hopkins et al. (2006)

#### hierarchical growth



Available supply of cold gas becomes exhausted

← exits loop... epoch of massive galaxy building and luminous quasars draws to a close





growth of rmassive black holes





Measuring M<sub>BH</sub>: (another talk)

 $M = k_1 \times v^2 R_{BLR} / G$ 

where v = line width  $R_{BLR} = k_2 \times L^{0.5}$ 

from 'simple'
photoionization physics
and measurements of
broad emission line
reverberation

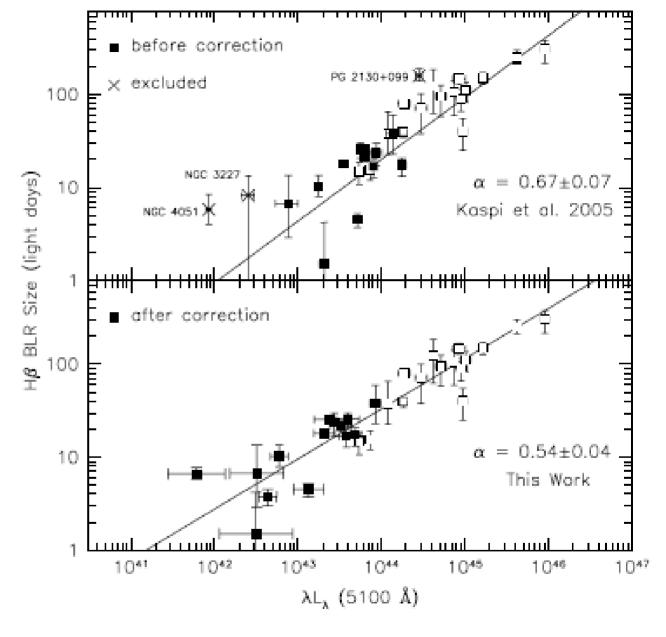
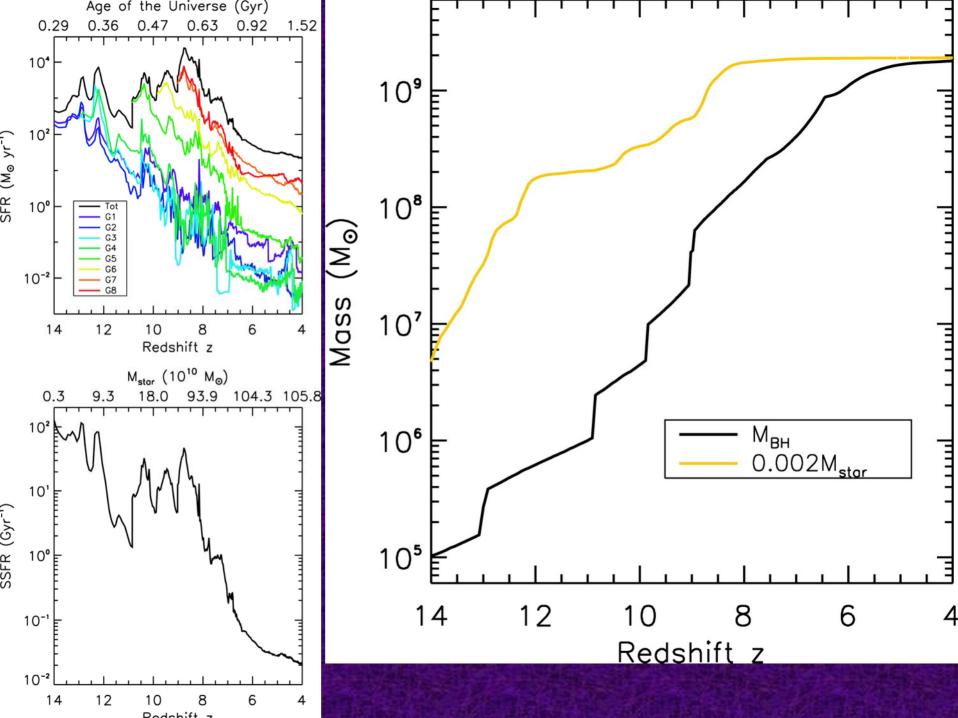
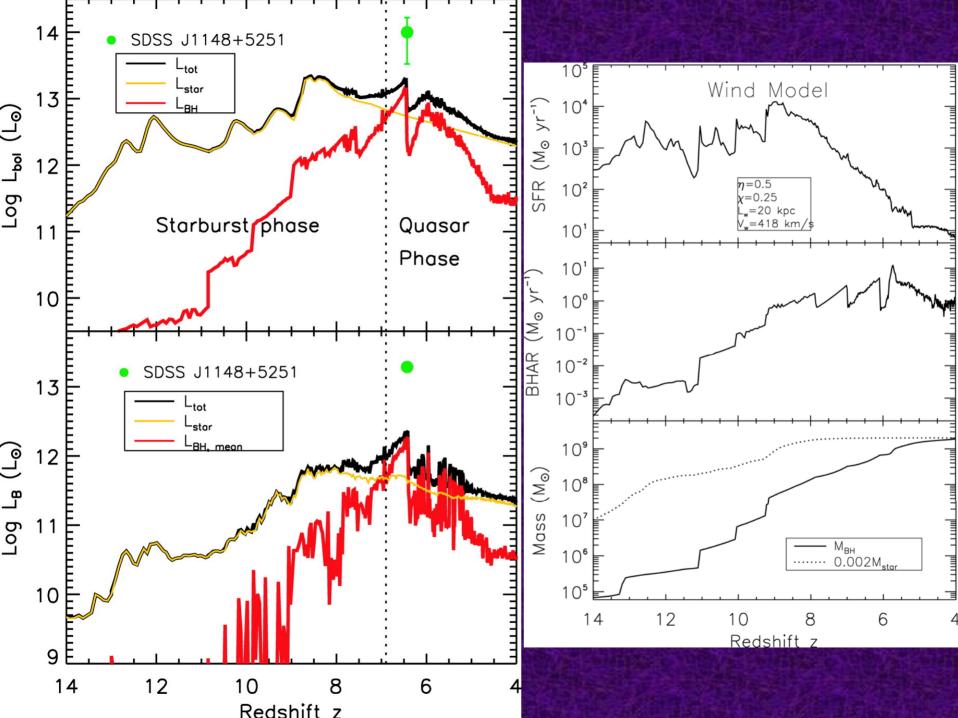


Figure 1. The top panel shows the recalibration of the R-L relationship from Kaspi et al. (2005) using the reanalyzed reverberation results of Peterson et al. (2004).





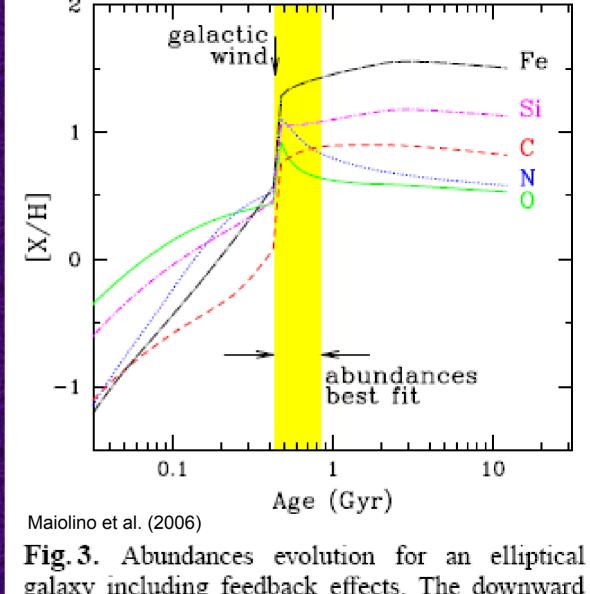


Fig. 3. Abundances evolution for an elliptical galaxy including feedback effects. The downward arrow indicates the onset of the galactic wind. The shaded area indicates the abundance sets which best fit the line ratios observed in the QSO spectra.

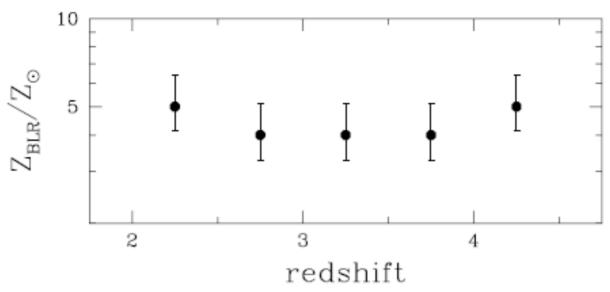


Fig. 38. Estimated metallicities from our composite spectra, averaged in the luminosity range  $-25.5 > M_B > -28.5$ , as a function of redshift. The estimation of the metallicity given in this figure is derived from the fit with the varying  $\beta$  and  $\Gamma$ , which are presented in Tables 12–16.

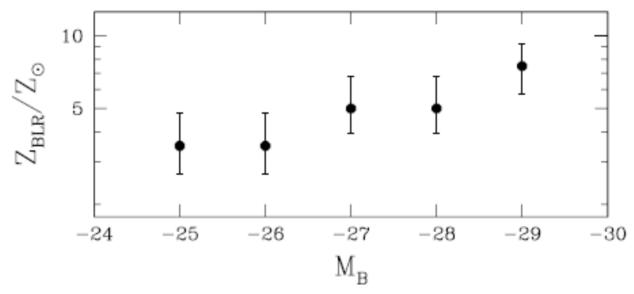


Fig. 39. Estimated metallicities from our composite spectra, averaged in the redshift range  $2.0 \le z < 3.0$ , as a function of luminosity. The estimation of the metallicity given in this figure is derived from the fit with the varying  $\beta$  and  $\Gamma$ , which are presented in Tables 12 and 13.

#### Centaurus A, 11 million ly away

a disturbed giant elliptical galaxy.

Deep imaging reveals rings of stars.



